



Session II. Engineering Collaboration Items

FPGA backend for a wideband receiver system

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SKA-JP Engineering Working Group

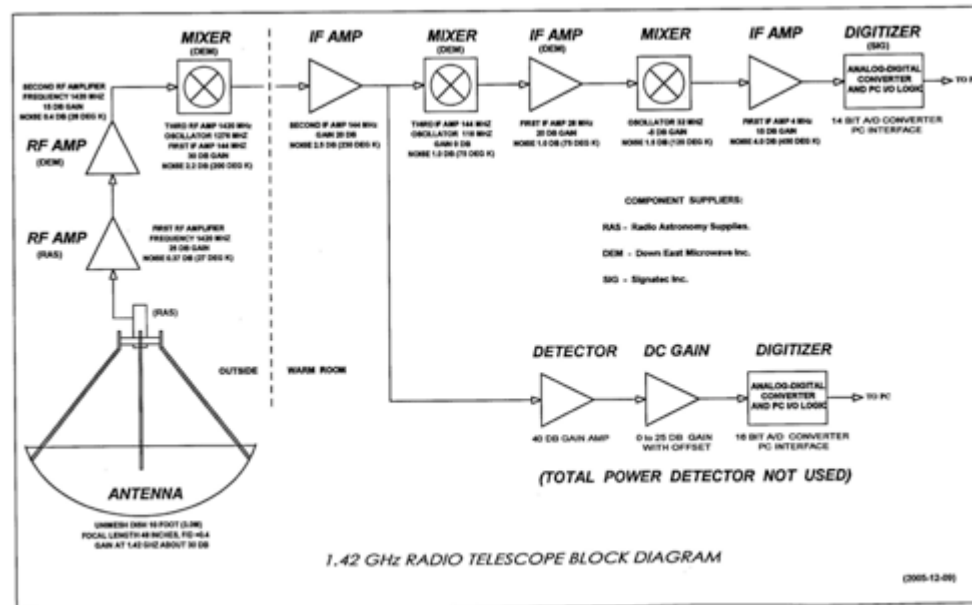
- SKA-JP EWG formed in 2012 (this year)
- Sub Working Groups
 - Digital Signal Processing (HN)
F-engine with ROACH
 - Software (Tomoharu Kurayama)
Tool for Faraday tomography
 - Feed for WBSPF (Hideki Ujihara)
Decade band feed

Primary Characteristics of SKA and Related Challenges

1. High Sensitivity : Large collecting area
→ Square Kilometer Array
2. High Resolution : Large physical distance
3. Wide Field : Small dish or Phased Array Field
4. Wide Band : Wideband feed and Fast A/D etc.

What's important for Wideband

1. Frequency response in analog components
2. Sampling rate at Analog-to-Digital converter

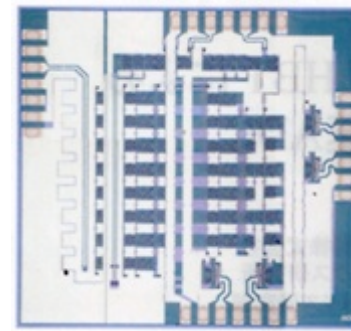
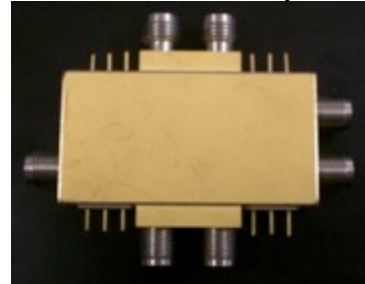
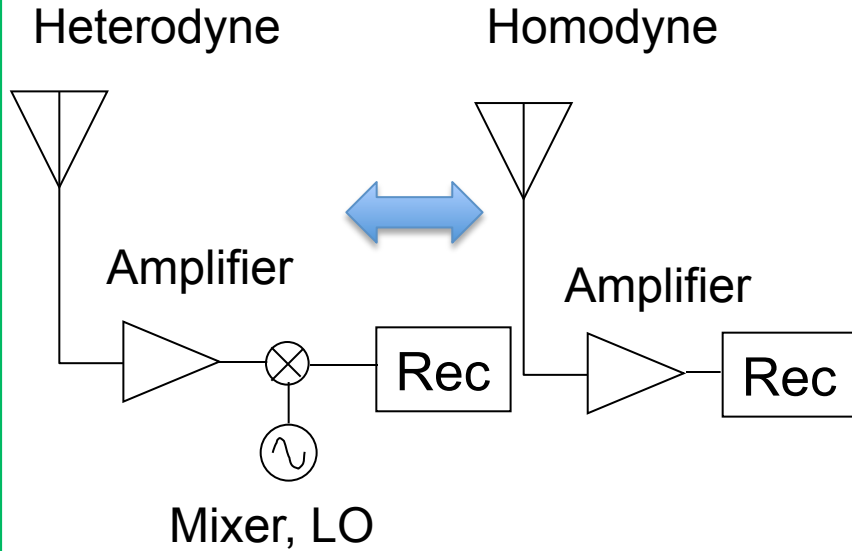


The 1.42 GHz radio telescope block diagram, courtesy of Fred Babott.

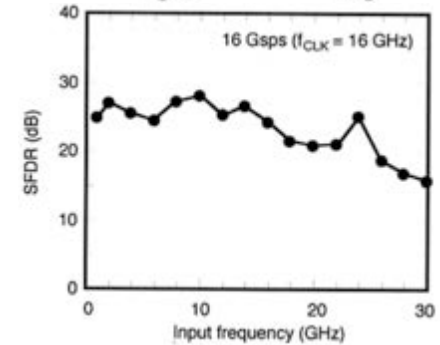
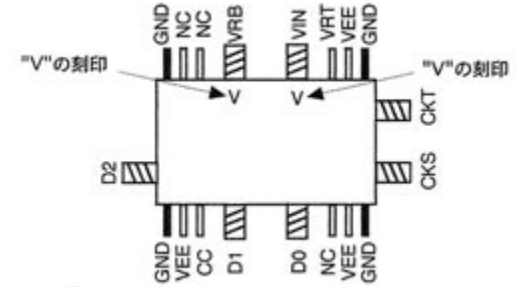
Less analog components and faster A/D

Homodyne receiver with Fast A/D converter

Fast InP HBC 3bit sampler (Kawaguchi@SKA-JP WS 2008)

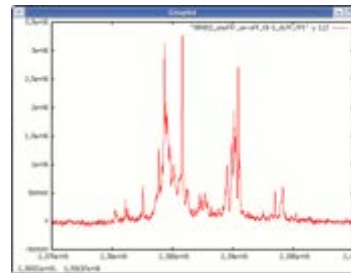
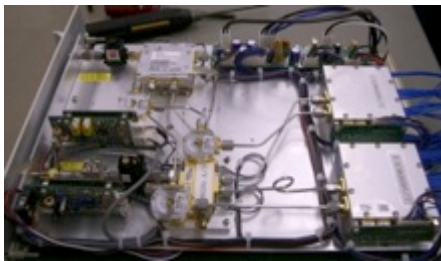


(3 mm × 3mm)

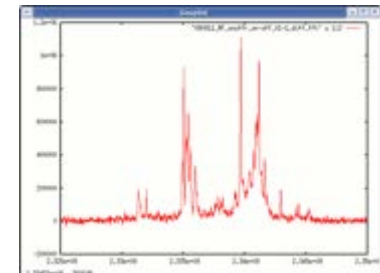


NTT Photonics Laboratory

Experiment



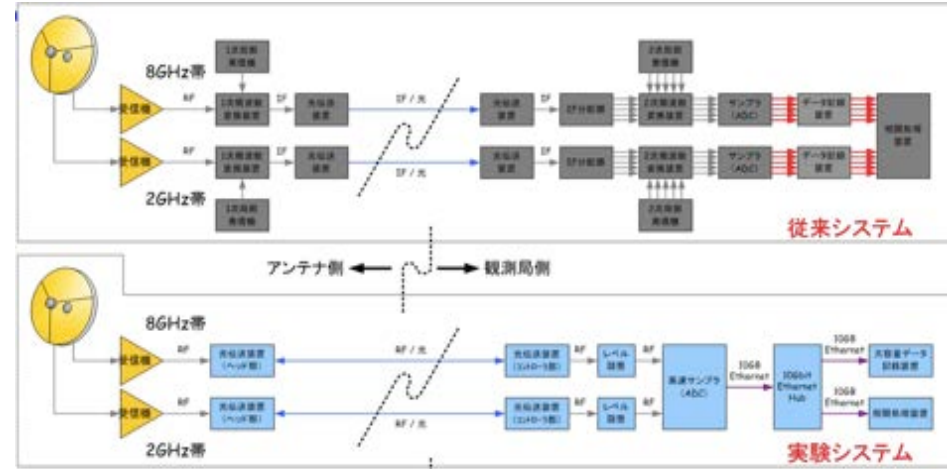
Spectra obtained with heterodyne receiver ($f_{LO}=(16.85+3)$ GHz, $f_{IF}=2.2-2.6$ GHz, $f_{sample}=8.192$ GHz)



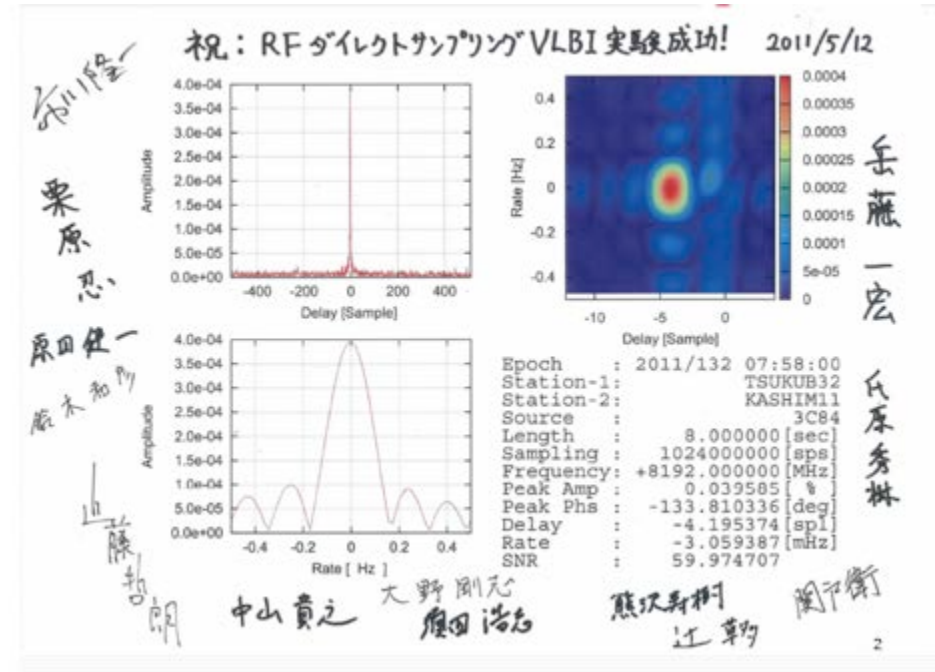
Spectrum obtained with homodyne receiver ($f_{RF}=20.480-24.576$ GHz, $f_{sample}=8.192$ GHz, $BW=4.096$ GHz)

VLBI experiments with Homodyne

Tsukuba 32m



Kashima 11m





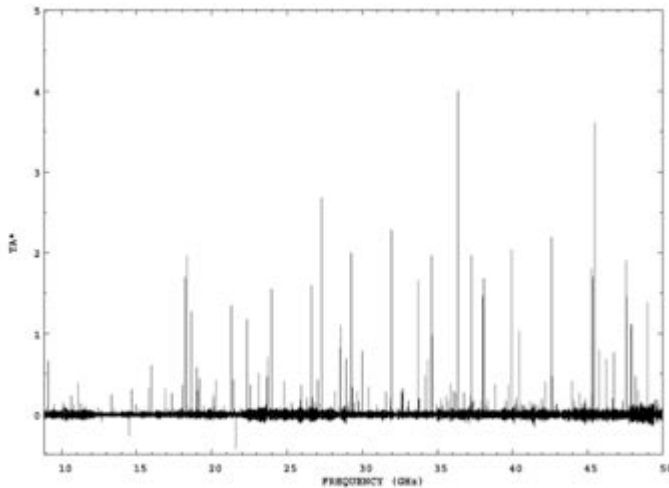
Homodyne Receiver System in the Future

- Experiments shows that homodyne receiver system is promising for the future radio telescope
- Wideband of Octave band (1:2) – Decade band (1:10) (Century band (1:100)?) would be common

Scientific Requirement

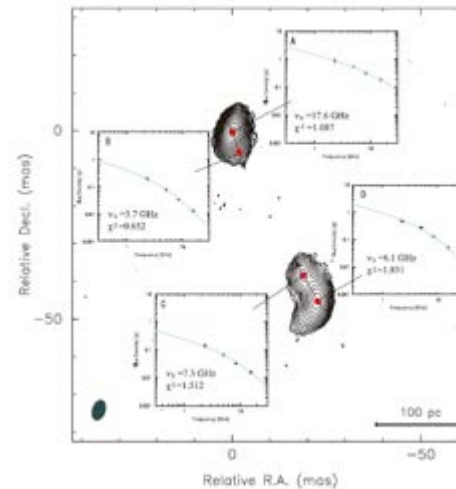
• Wideband spectrum

Line observation



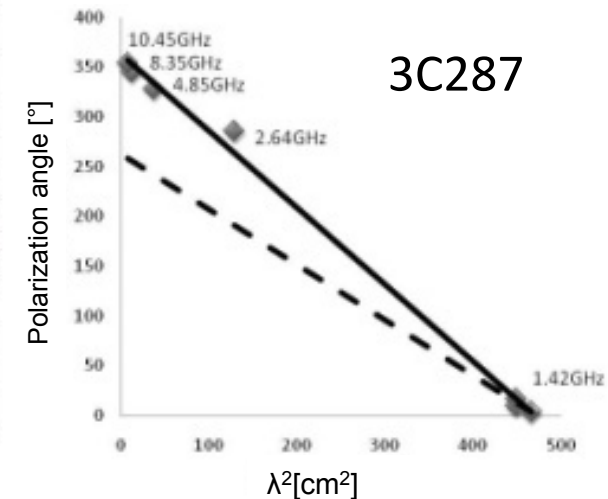
Ultra wideband Line survey for TMC-1 in 8-50GHz (Kaifu et al. 2004)

Cont. observation



Wideband SED (2.2-15.3GHz) of AGN CTD93 (Nagai et al. 2006)

Polarization observation



Polarization angle over wideband → More accurate rotation measure (slope)

• Higher sensitivity: $\Delta T \propto B^{-1/2}$

- Distant faint object can be detected
- Increase of reference source for astrometry



Wideband would be great but ...

Another issue shows up:

Bandwidth ratio to observed frequency is not negligible anymore.

For example:

The case of multi-line observation:

$$\text{HI: } f_0 = 1.4\text{GHz}, \Delta f = 48\text{kHz} \rightarrow \Delta v = \Delta f / f_0 = 10\text{km/s}$$

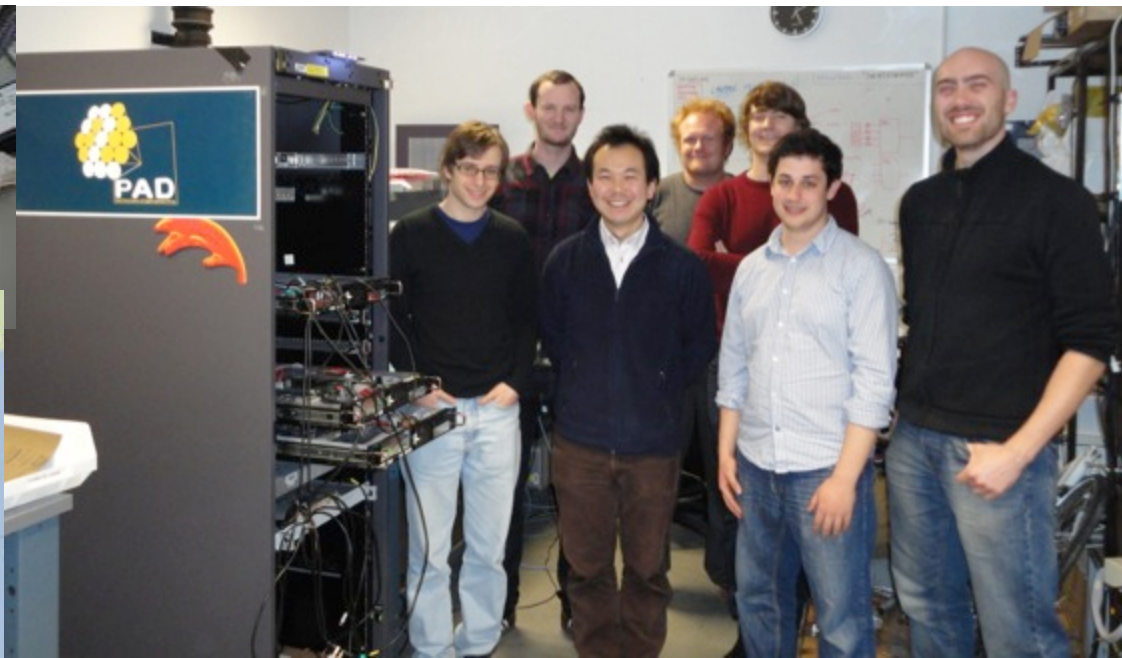
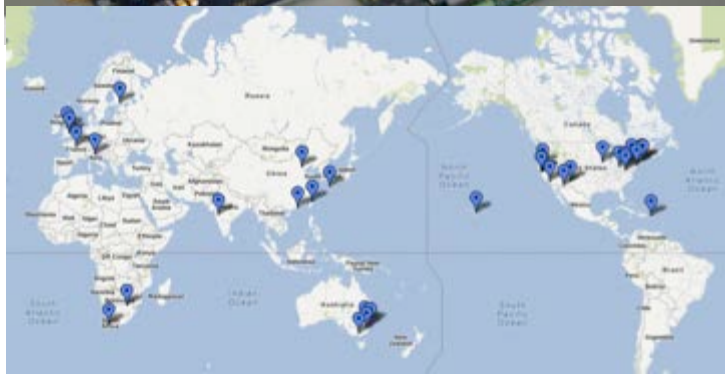
$$\text{CH}_3\text{OH: } f_0 = 6.7\text{GHz}, \Delta f = 48\text{kHz} \rightarrow \Delta v = \Delta f / f_0 = 2\text{km/s}$$

If frequency coverage is 1-8GHz, 150000ch is necessary (required velocity resolution: 10km/s), but it's oversampling for the high frequency.

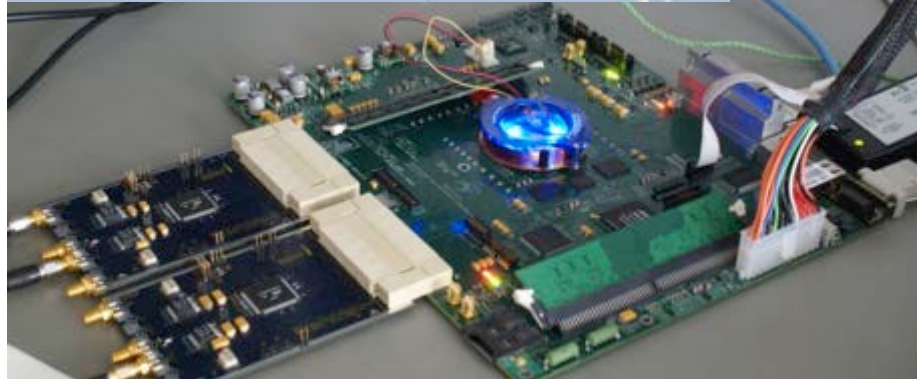
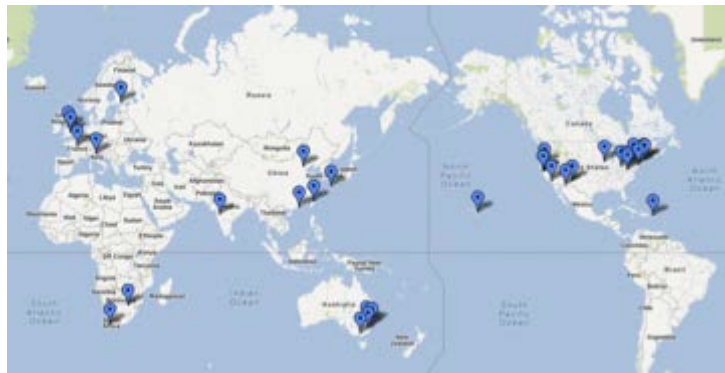
Optimized channelization

Our current interest

- Optimized channelization (ex. $\Delta f/f_0 = \text{const.}$) in F-engine using ROACH board (Collaboration with Oxford)
- 2 ROACH boards were recently delivered to Kagoshima university.



What is ROACH



CASPER
COLLABORATION FOR ASTRONOMY SIGNAL PROCESSING AND ELECTRONICS RESEARCH

Home Group Documentation Mail Archive About

Goals

The primary goal of CASPER is to streamline and simplify the design flow of radio astronomy instrumentation by promoting design reuse through the development of platform-independent, open-source hardware and software.

Our aim is to couple the real-time streaming performance of application-specific hardware with the design simplicity of general-purpose software. By providing parameterized, platform-independent "gateway" libraries that run on reconfigurable, modular hardware building blocks, we abstract away low-level implementation details and rapidly design and deploy new instruments.

News

Journal of Astronomical Instrumentation
January 26, 2012
[New opportunities for publication in Astronomical Instrumentation](#) (Lincoln Greenhill)

CASPER Workshop 2011
August 1, 2011
[CASPER 2011 Workshop website is now up and ready for registration](#) (Yashwant Gupta)

CASPER Workshop 2011
May 25, 2011
[First announcement of the CASPER 2011 Workshop](#) (Yashwant Gupta)

New Memo
November 23, 2010
[Optimize CASPER Development by "Black Boxing" Designs](#)
This memo describes how to use System Generator's Black Box to develop System Generator models more efficiently. (David MacMahon)

Workshop program now available.
July 22, 2010
[Casper Workshop 2010 Program](#) (Jonathan Weintraub)

Casper Workshop 2010 Announced!
March 16, 2010
[Online registration is now open.](#)

- CASPER is a project to develop open-source hardware and software.
 - Good starting point for international collaboration
- ROACH (**R**econfigurable **O**pen **A**rchitecture **C**omputing **H**ardware) is an FPGA board developed by CASPER project



Comparison of Rival Architectures

FPGA (ROACH)	GPU (Tesla C2070)	CPU (i7 3960X)
High I/O (GbE, InfiniBand)	Low I/O (PCIe)	Low I/O
Low power consumption (15 W)	High power consumption (225W)	High power consumption (130 W)
Difficult to program (HDL)	Easy to program (CUDA, OpenCL)	Easy to program (multiple languages)

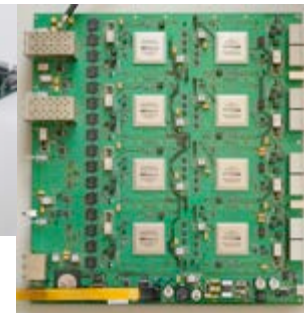
Slide from Chris Taylor, "Thoughts on CASPER 2011", University of Manchester GPU Club

- ASIC has the lowest power consumption and good I/O as FPGA. Quite difficult to develop.

ROACH vs UniBoard

- ROACH and UniBoard have the same philosophy:
 “open-source” library
- UniBoard has the better cost and power performance
- ROACH seems still easier to use
 ROACH : MATLAB/SIMULINK
 UniBoard: HDL
- ROACH is easier to start
 We tried to purchase UniBoard last year but was not ready
- CASPER international team is well organized

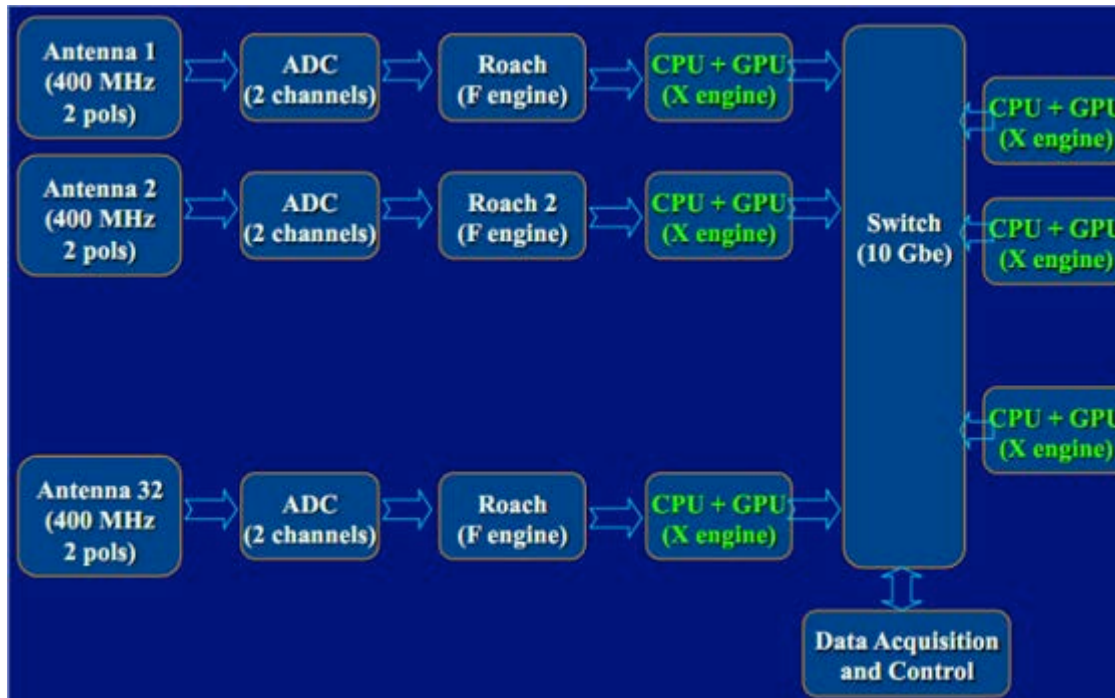
	ROACH	UniBoard
FPGA chip	1 x Xilinx Virtex 5	8 x Altera Stratix IV
Leading Institute	UC Berkley	ASTRON
Cost / chip	\$5,500	\$2,700



It seems good to start DSP development with ROACH and copy to UniBoard in the future

Possible EA collaboration

- Fast A/DC board (Taiwan ?)
First step: just use ADC1x5000-8, Second step: development of faster A/D ?
- Hybrid architecture (F-engine with FPGA and X-engine with GPU) (Korea ?)
- And more ...



- ✧ F Engine:
 - ✓ FFT
 - ✓ Channelise
 - ✓ Packet data
- ✧ X Engine:
 - ✓ Correlate
 - ✓ Integrate
 - ✓ Dump data



Summary

- SKA-JP EWG formed this year
- The keyword is “Wideband” and fast A/D is the key component.
- The homodyne receiver system might be common technique in the near future
- SKA-JP DSP team started developing F-engine using ROACH board.
- Collaboration related to A/D and hybrid-architecture correlator.